

# Physical interpretation of the Weert superpotential

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**Abstract.** Weert [1] obtained a potential for the bounded part of the Maxwell Tensor which is associated to the Liénard-Wiechert field. We show that this potential can be interpreted as an intrinsic Angular Momentum Density for the corresponding Electromagnetic Field.

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## 1. Introduction

In this paper we deal with electromagnetic field which is produced by a point particle charge in the Minkowski space. This charge gives the Liénard-Wiechert field (LW) from which the Maxwell tensor  $T_{jc}$  can be obtained. The last can be separated in two parts:  $T_{jc}^B$  the bounded one, and  $T_{jc}^R$  the radiative part (in the sense of Teitelboim [2]).

Sec. 2 is devoted to a brief exposition of the LW field. Weert [1] constructed a potential  $K_{jbc}^B$  for  $T_{jc}^B$ , this will serve us to propose, in Sec. 3, a physical interpretation of this potential using the Bhabha [4]-Synge [5] region; moreover, if we decompose  $K_{jbc}^B$  in two parts which satisfy the symmetry of the Lanczos Spintensor [6], we obtain the rupture of  $T_{jc}^B$  proposed by López [9], which is very important in the study of the angular momentum of the LW field. In Sec. 4 we construct a non-local Superpotential (depending of the past history of the charge) for the radiative part; and we show the terms of  $T_{jc}^R$  which do not participate in the flux of energy and momentum through the Bhabha-Synge tube.

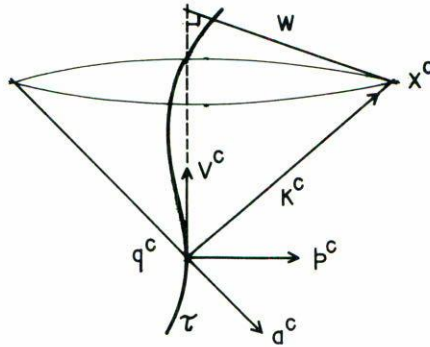


FIGURE 1. Kinematics of the world-line.

2. Point charge in arbitrary motion

A charged particle moving in an arbitrary motion in the Minkowskian space, produced the electromagnetic LW field with the 4-potential and Faraday tensor given by

$$A^c = qw^{-1}v^c, \quad F_{bc} = qw^{-2}(U_bK_c - U_cK_b), \tag{1.a}$$

and the Maxwell Tensor

$$T_{bc} = q^2w^{-4} \left[ K_bU_c + K_cU_b + (a^2 - B^2)K_bK_c + \frac{1}{2}g_{bc} \right], \tag{1.b}$$

where (Fig. 1):

$$(x^c) = (x, y, z, t), \quad (g_{bc}) = \text{Diag}(1, 1, 1, -1)$$

$\tau$  is the proper timer,  $q^c(T)$  will be the retarded point,  $v^c$  is the 4-velocity,  $a^c$  the 4-acceleration

$$K^c = x^c - q^c \tag{1.c}$$

$w = -K^c v_c$  which is the so called retarded distance

$$W = -K^c a_c, \quad a^2 = a^c a_c$$

$B = w^{-1}(1 - W)$  known as the Plebański [10] invariant

$$U_c = Bv_c + a_c, \quad p^c = w^{-1}K^c - v^c.$$

Teitelboim [2] proved that Eq. (1.b) can be written as the sum of a bounded part  $T_{bc}$  and a radiative one  $T_{bc}$ , namely,

$$T_{bc} = T_{B\ bc} + T_{R\ bc}, \tag{2.a}$$

such that

$$T_{B\ bc} = q^2 w^{-4} \left[ \frac{1}{2} g_{bc} + (K_b a_c + K_c a_b) + B(K_b v_c + K_c v_b) - w^{-2}(1 - 2W)K_b K_c \right], \tag{2.b}$$

and

$$T_{R\ bc} = q^2 w^{-2}(a^2 - w^{-2}W^2)K_b K_c. \tag{2.c}$$

Each tensor possesses the following differential properties

$$T_{B\ b\ ,c}^c = 0, \tag{2.d}$$

$$T_{R\ b\ ,c}^c = 0, \tag{2.e}$$

respectively, when it is valued out of the world-line.

Weert [1] proved that Ec. (2.d) comes out from the existence of the superpotential

$$K_{B\ jbc} = -q^2/4w^{-4} \left[ w^{-1}(3 - 4W)(v_j \times K_b)K_c + 4(a_j \times K_b)K_c + g_{cj}K_b - g_{cb}K_j \right], \tag{3.a}$$

where we have used the Lowry [3] notation

$$A_j \times B_c = A_j B_c - A_c B_j, \tag{3.b}$$

such that

$$T_{B\ bc} = K_{B\ b\ c,j}^j. \tag{3.c}$$

From here, (2.d) follows immediately.

In the next section we will study (3.a) in order to give a physical interpretation of this superpotential. In Sec. 4 we will do a brief analysis of Eqs. (2.c-e).

### 3. Weert superpotential

Our point charge in arbitrary motion gives an electromagnetic field which possesses an intrinsic angular momentum (IAM). Here we will prove that  $K_{B\ jbc}$  behaves like a

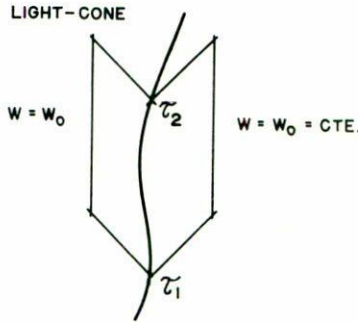


FIGURE 2. Bhabha-Syngé tube.

density for IAM when the corresponding fluxes are calculated through a Bhabha [4]-Syngé [5] tube.

The superpotential (3.a) has the following properties

$$\begin{aligned}
 K_B{}_{jbc} &= -K_B{}_{bjc} && \text{antisymmetry,} \\
 K_B{}^c{}_j &= 0 && \text{null trace,} \\
 K_B{}_{jbc} + K_B{}_{bcj} + K_B{}_{cjb} &= 0 && \text{cyclic,} \\
 K_B{}^c{}_{jbc} &= 0 && \text{null divergence.}
 \end{aligned}
 \tag{4}$$

These properties agree with those of the Lanczos [6] Generator  $K_{jbc}$  for the Weyl tensor of the space-time. Lanczos calculated  $K_{jbc}$  for weak gravitation fields and in his analysis, the Dirac equation, for spin  $\frac{1}{2}$ , appeared. For this reason he called the potential  $K_{jbc}$  spin-tensor. In our case,  $K_B{}^c{}_{jbc}$  will be associated with the IAM of the LW field.

Consider the Bhabha-Syngé tube (Fig. 2), which is composed by the light cones with the tops in  $\tau = \tau_1$  and  $\tau = \tau_2$ , and a surface of constant retarded distance. First, let us calculate the  $K_B{}_{jbc}$  flux through a light cone; the expression is given by Syngé [5]

$$\dot{M}_{jb} = \int_{\tau=\text{const}} K_B{}_{jbc} d\sigma^c = - \int_0^{w_0} w dw \int_{\tau=\text{const}} d\Omega K_B{}_{jbc} K^c, \tag{5.a}$$

where  $d\Omega$  is the element of solid angle.

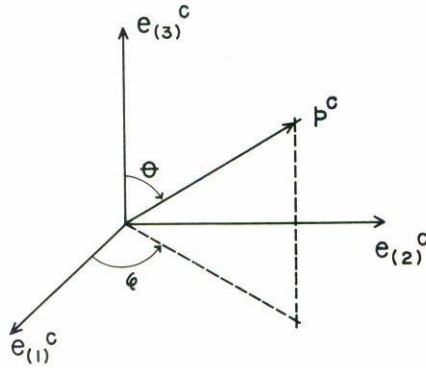


FIGURE 3. Fermi tetrad.

The unitary space-like vector  $p^c$  has been defined in Fig. 1

$$p^c = w^{-1}k^c - v^c, \quad p^c p_c = 1, \quad p^c v_c = 0. \tag{5.b}$$

Now for any event in the line-universe,  $p^c$  can be written in terms of a Fermi tetrad  $e_{(\gamma)c}$  with  $\gamma = 1, 2, 3$  (Fig. 3).

$$\begin{aligned} \frac{de_{(\gamma)}^c}{d\tau} &= a_{(\sigma)}v^c = (a^r e_{(\sigma)r})v^c, \\ p^c &= \sin\theta \cos\phi e_{(1)}^c + \sin\theta \sin\phi e_{(2)}^c + \cos\theta e_{(3)}^c \\ &= P_{(\sigma)}e_{(\sigma)}^c = (P^r e_{(\sigma)r})e_{(\sigma)}^c \end{aligned} \tag{5.c}$$

and  $d\Omega = \sin\theta d\theta d\phi$ .

From (3.a), it is clear that

$$K_{\text{B}}{}^{jbc}K^c = 0. \tag{5.d}$$

Therefore (5.a) implies  $\tilde{M}_{jb} = 0$ , that is,  $K_{\text{B}}{}^{jbc}$  flux vanishes through a light cone.

Now we will calculate the  $K_{\text{B}}{}^{jbc}$  flux through the 3-space  $w = w_0 = \text{const}$ ; the expression is given in Sygne [5]

$$M_{\text{I}}{}^{jb} = \int_{w=w_0} K_{\text{B}}{}^{jbc} d\sigma^c = w^2 \int_{\tau_1}^{\tau_2} d\tau \int K_{\text{B}}{}^{jb}{}^c w_{,c} d\Omega, \tag{6.a}$$

where

$$w_{,c} = \text{Gradient of } w = -v_c + BK_c \tag{6.b}$$

Therefore, Ecs. (3.a, 5.c, 6.a, b) imply that

$$M_{I\ jb} = \frac{8\pi}{3} q^2 \int_{\tau_1}^{\tau_2} (v_j \times a_b) d\tau. \tag{6.c}$$

This last result agrees with the intrinsic angular momentum of the LW Field [7]. Therefore,  $K_{B\ jbc}$  behaves as a density for such an angular momentum.

The superpotential (3.a) accepts the following rapture

$$K_{B\ jbc} = \tilde{K}_{B\ jbc} + \bar{K}_{B\ jbc}, \tag{7.a}$$

where

$$\tilde{K}_{B\ jbc} = q^2 w^{-4} [(-a_j + w^{-1} W v_j) \times K_b] K_c \tag{7.b}$$

and

$$\bar{K}_{B\ jbc} = -w^{-4} [g_{cj} K_b - g_{cb} K_j + 3w^{-1} (v_j \times K_b) K_c]. \tag{7.c}$$

The potentials in Eqs. (7.b, c) satisfy all the properties in (4), as the Lanczos Spin-tensor does.

It is simple to prove that

$$\int_{\tau=\text{const}} \bar{K}_{B\ jbc} d\sigma^c = \int_{w=\text{const}} \bar{K}_{B\ jbc} d\sigma^c = 0, \tag{8.a}$$

and

$$M_{I\ jb} = \int_{w=\text{const}} \tilde{K}_{B\ jbc} d\sigma^c \quad \text{and} \quad \int_{\tau=\text{const}} \tilde{K}_{B\ jbc} d\sigma^c = 0. \tag{8.b}$$

That is,  $\bar{K}_{B\ jbc}$  doesn't contribute to the IAM of the electromagnetic field; this means that  $\tilde{K}_{B\ jbc}$  is the active part of part of  $K_{B\ jbc}$ .

The result can be obtained by using Stokes Theorem and the Rowe [8] identity

$$\bar{K}_{B\ jbc} = \left( \frac{q^2}{4} w^{-4} D_{jbc} \tau \right)_{,r}, \tag{8.c}$$

where  $D_{jbc\tau}$  is a tensor used by Synge [5] in another context

$$D_{sarb} = (g_{sr}K_b - g_{sb}K_r)K_a + (g_{ab}K_r - g_{ar}K_b)K_s. \tag{8.d}$$

Finally using (7.a) in (3.c), the following decomposition is obtained

$$T_{B\ bc} = \tilde{T}_{B\ bc} + \bar{T}_{B\ bc}, \tag{9}$$

with  $\tilde{T}_{B\ bc} = \hat{K}_{B\ b\ c, j}^j$  and  $\bar{T}_{B\ bc} = \bar{K}_{B\ b\ c, j}^j$ .

Eq. (9) is important at the moment we relate it to the angular momentum radiated by the charge [9].

#### 4. On a rupture for $T_{R\ ab}$

Here we show how the radiative part of  $T_{ab}$  can be written as the sum of two terms; one of them doesn't participate in the energy and momentum flux through the Bhabha-Syngé tube. Moreover, we will give a potential for  $T_{R\ bc}$ .

The expression (2.c) can be written in the form

$$T_{R\ bc} = T_{i\ bc} + \tilde{T}_{R\ bc}, \tag{10.a}$$

where

$$T_{i\ bc} = q^2 w^{-4} (a^2 - 3w^{-2}W^2) K_b K_c, \tag{10.b}$$

$$T_{i\ b\ .c} = 0$$

$$T_{R\ bc} = 2q^2 w^{-6} W^2 K_b K_c, \tag{10.c}$$

$$T_{R\ b\ .c} = 0$$

It is simple to demonstrate that

$$\int_{\substack{\tau = \text{const} \\ \text{or} \\ w = \text{const}}} T_{i\ bc} d\sigma^c = 0, \tag{11.a}$$

that is, (10.b) doesn't contribute to the energy and momentum flux through the Bhabha-Syngé tube. In a similar way

$$\int_{\substack{w = \text{const} \\ \text{or} \\ \tau = \text{const}}} (x^j T_{i\ bc} - x^b T_{i\ jc}) d\sigma_c = 0. \tag{11.b}$$

Hence  $T_{bc}$  doesn't participate either in the momentum fluxes for such tube. Due to Ecs. (11.a, b) we say that the tensors in (10.b) represent the inactive part of  $T_{ab}$  with respect to the Bhabha-Syngé region.

The conservation law is immediately deduced from the existence of the superpotential:

$$K_{jbc} = -\frac{q^2}{4}w^{-2} \left[ w^{-2}W^2(g_{cj}K_b - g_{cb}K_j) + w^{-1}W(v_j \times K_b)(a_c - 3w^{-2}WK_c) + (a_j \times K_b)(4w^{-2}WK_c - a_c) \right], \tag{12.a}$$

such that

$$T_{bc} = K_{b \ c, j}^j. \tag{12.b}$$

Moreover, the identity  $T_{R \ b, c}^c = 0$  is a consequence of

$$\tilde{T}_{R \ bc} = \tilde{K}_{R \ b \ c, j}^j, \tag{13.a}$$

where

$$\tilde{K}_{R \ bjc} = -2qF_{bj}p(\sigma)p(\gamma) \left[ \int_0^\tau a(\sigma)a(\gamma)v_c d\tau + p(\beta) \int_0^\tau a(\sigma)a(\gamma)e_{(\beta)c} d\tau \right] \tag{13.b}$$

in equation (13.b), the sum over  $\sigma, \gamma, \beta = 1, 2, 3$  has to be done.

To verify Eq. (13.a), we have to use the following relations

$$\begin{aligned} \tau_{,j} &= -w^{-1}K_j, && \text{retarded derivative} \\ F_{b \ ,j}^j &= 0, && \text{Maxwell equations} \\ F_{b \ ,j}^j &= qw^{-2}K_b, && \text{null eigenvector} \\ F_b^j p(\sigma)_{,j} &= 0, && \text{Fermi tetrad} \\ W &= -wp(\sigma)a(\sigma), \\ k^c &= w(v^c + p^c). \end{aligned} \tag{13.c}$$

The existence of integrals in equation (13.b) shows the non-local character of  $\tilde{K}_{R \ bjc}$ ; that means that it depends on the past history of the charge. Therefore



Eqs. (2.a, 3.c, 10.a, 12.b, 13.a) imply

$$T_{bc} = \left( K_B^j{}_c + K_i^j{}_c + \tilde{K}_R^j{}_c \right)_{,j}. \quad (14)$$

Hence the Maxwell tensor associated to the LW Field is an *exact divergence*.

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**Resumen.** Weert [1] obtuvo un potencial para la parte acotada del tensor de Maxwell asociado al campo de Liénard-Wiechert. Aquí mostramos que dicho potencial puede interpretarse como una densidad de momento angular intrínseco del correspondiente campo electromagnético.